TRAINING

Everything must be fully justified on the basis of the definitions or the appendix.

Problem Cosmology

The Friedmann–Lemaître–Robertson–Walker metric describes a homogeneous, isotropic, expanding or contracting universe:

$$ds^{2} = c^{2} dt^{2} - a(t)^{2} \left(\frac{dr^{2}}{1 - kr^{2}} + r^{2} d\Omega^{2} \right) \text{ with } d\Omega^{2} = d\theta^{2} + \sin^{2}\theta d\phi^{2} \text{ and } (x^{0} = ct, x^{1} = r, x^{2} = \theta, x^{3} = \phi).$$

k is a constant representing the curvature of the space. k belong to the set $\{-1, 0, +1\}$. For k=-1, space has an open hyperbolic curvature, and after the Big Bang the universe will expand indefinitely. For k=1, space has a closed elliptic curvature, and after expansion, contraction will follow until a Big Crunch. Let's consider the critical case k=0 and the current scale factor $a(t)=e^{H_0t}$ with H_0 the Hubble constant.

- 1) Calculate the components Γ^0_{11} , Γ^2_{21} and Γ^1_{23} of the Christoffel symbols.
- 2) Express the component R^{1}_{212} of the Riemann curvature tensor for this metric. Is spacetime flat?
- 3) Then, find the Ricci tensor R_{00} component, and the Ricci scalar R.
- 4) Using Einstein's equation find the expression for the critical matter density ρ_c of the universe according to FLRW's model. Isolated particles are considered: $T^{\mu\nu} = \rho_0 u^{\mu} u^{\nu}$.
- 5) Let's consider the spatial part dl^2 of the metric. Find the space Riemann curvature tensor components. Is space flat?
- 6) Give an estimation of the numerical value for the critical density ρ_c with $H_0 \simeq 2 \times 10^{-18} \, s^{-1}$ and $G \simeq 7 \times 10^{-11} \, m^3 \, s^{-2} \, kg^{-1}$. What future for the universe with the experimental value $\rho \simeq 7 \times 10^{-30} \, kg/m^3$?

APPENDIX:

Connections (Christoffel symbols): $\Gamma^{\alpha}_{\ \mu\nu} = \frac{1}{2} g^{\alpha\beta} (\partial_{\mu} g_{\beta\nu} + \partial_{\nu} g_{\beta\mu} - \partial_{\beta} g_{\mu\nu})$

Results given (non null connections): $\Gamma^{1}_{10} = \Gamma^{2}_{20} = \Gamma^{3}_{30} = H_{0}/c$, $\Gamma^{3}_{31} = \Gamma^{2}_{21}$, $\Gamma^{0}_{22} = \frac{H_{0}}{c}r^{2}e^{2H_{0}t}$,

$$\Gamma^{1}_{22} = -r, \quad \Gamma^{0}_{33} = \frac{H_{0}}{c} r^{2} \sin^{2}\theta \, e^{2H_{0}t}, \quad \Gamma^{1}_{33} = -r \sin^{2}\theta, \quad \Gamma^{2}_{33} = -\sin\theta \cos\theta \quad and \quad \Gamma^{3}_{32} = \cos\theta/\sin\theta.$$

Riemann tensor: $R^{\alpha}_{\beta\gamma\delta} = \Gamma^{\alpha}_{\beta\delta,\gamma} - \Gamma^{\alpha}_{\beta\gamma,\delta} + \Gamma^{\alpha}_{\sigma\gamma}\Gamma^{\sigma}_{\beta\delta} - \Gamma^{\alpha}_{\sigma\delta}\Gamma^{\sigma}_{\beta\gamma}$

Results given (non null components): $R^0_{202} = r^2 R^0_{202} = R^1_{212}$ and $R^0_{303} = R^1_{313} = R^2_{323} = \sin^2 \theta R^1_{212}$.

Ricci tensor: $R_{\mu\nu} = R^{\alpha}_{\mu\nu\alpha}$ Ricci scalar: $R = g^{\mu\nu}R_{\mu\nu}$

Results given: $R_{11} = 3 \frac{H_0^2}{c^2} e^{2H_0 t}$, $R_{22} = r^2 R_{11}$ and $R_{33} = \sin^2 \theta R_{22}$.

Einstein's Equation: $R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = \frac{8\pi G}{c^4} T_{\mu\nu}$

Page 1/1 Mathieu Rouaud